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Abstract

We investigate the structure of a Finsler manifold of nonnegative weighted Ricci curvature including a straight line, and extend the classical Cheeger-Gromoll-Lichnerowicz splitting theorem. Such a space admits a diffeomorphic, measure-preserving splitting in general. As for a special class of Berwald spaces, we can perform the isometric splitting in the sense that there is a one-parameter family of isometries generated from the gradient vector field of the Busemann function. A Betti number estimate is also given for Berwald spaces.

1 Introduction

The Ricci curvature is one of the most important quantities in geometry and analysis on Riemannian manifolds. The Ricci curvature (or tensor) plays prominent roles in various ways, from the classical comparison theorems due to Rauch and Bishop to Hamilton and Perelman's celebrated theory of the Ricci flow. Recently, it turned out that the Ricci curvature is quite useful also in the study of Finsler manifolds. A Finsler manifold is a manifold endowed with a (Minkowski) norm on each tangent space. Inspired by the theory of weighted Riemannian manifolds, the *weighted Ricci curvature* Ric_N was introduced in [Oh3] for a Finsler manifold (M, F) equipped with an arbitrary measure m on M , where $N \in [\dim M, \infty]$ is a parameter (see Definition 2.4). Bounding Ric_N from below by $K \in \mathbb{R}$ (i.e., $\text{Ric}_N(v) \geq KF(v)^2$) is equivalent to Lott, Sturm and Villani's *curvature-dimension condition* $\text{CD}(K, N)$ ([Oh3]). This equivalence has many applications via the general theory of the curvature-dimension condition, such as the Bishop-Gromov volume comparison and the Lichnerowicz inequality on the spectral gap (see [Oh3]). Furthermore, the Laplacian comparison theorem for a natural nonlinear Laplacian ([OS1]) as well as the Bochner-Weitzenböck formula ([OS3]) hold.

The aim of this article is to generalize another fundamental theorem in comparison geometry involving the Ricci curvature, Cheeger and Gromoll's *splitting theorem* ([CG1]), that asserts that a Riemannian manifold (M, g) of nonnegative Ricci curvature including

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a straight line admits an isometric splitting $M = M' \times \mathbb{R}$. This splitting theorem was extended to weighted Riemannian manifolds by Lichnerowicz and others ([Li], [FLZ], [WW]), and to (weighted) Alexandrov spaces by introducing appropriate notions of the lower Ricci curvature bound ([KS], [ZZ]).

In the Finsler case, as normed spaces (equipped with the Lebesgue measure) have the nonnegative Ricci curvature, the isometric splitting can not be expected. Nevertheless, our first main result (Corollary 4.4) asserts that a diffeomorphic and measure-preserving splitting holds for general Finsler manifolds. The proof is essentially parallel to the Riemannian case, thanks to the Laplacian comparison theorem in [OS1]. One can describe the splitting in more details for Berwald spaces. Roughly speaking, a Berwald space is a Finsler manifold modeled by a single normed space (for example, Riemannian manifolds and normed spaces are Berwald spaces). In the Berwald case, we can also split the metric in the sense that there exists an l -parameter family of isometries $\varphi_p : M \rightarrow M$, $p \in \mathbb{R}^l$, such that $\bigsqcup_{p \in \mathbb{R}^l} \varphi_p(\overline{M}) = M$, where the $(\dim M - l)$ -dimensional submanifold $\overline{M} \subset M$ is characterized by the property that any Busemann function is constant on \overline{M} (see Theorem 5.4 for the precise statement). A Betti number estimate along the lines of [CG1], [CG2] is also given in the Berwald case (Theorem 5.6).

The study of Finsler manifolds has an extra importance from the view of the curvature-dimension condition. As we mentioned above, Finsler manifolds give a nice class of model spaces satisfying the curvature-dimension condition. One can use this class as a test to see what properties of Riemannian manifolds of Ricci curvature bounded below can be or can not be expected to hold for general metric measure spaces satisfying the curvature-dimension condition (in other words, to see either such a property depends only on the ‘Ricci curvature bound’, or it also requires that the space is ‘Riemannian’). For instance, the contraction property of the heat flow with respect to the L^2 -Wasserstein distance fails on (non-Riemannian) Finsler manifolds ([OS2]). Recently, the contraction property was shown in [AGS] for metric measure spaces by assuming $\text{CD}(K, \infty)$ and the linearity of the heat flow (the linearity means that the space is ‘Riemannian’ in a sense). The Bochner-Weitzenböck formula also holds in such a case (see also [GKO]), whereas it is unclear how to remove the linearity.

The article is organized as follows. We review necessary notions in geometry and analysis on Finsler manifolds in Section 2. Section 3 is devoted to the study of Busemann functions. We show splitting theorems in Sections 4, 5 in the general and Berwald cases, respectively.

2 Geometry and analysis on Finsler manifolds

We review the basics of Finsler geometry (we refer to [BCS] and [Sh1] for further reading), and introduce the weighted Ricci curvature and the nonlinear Laplacian studied in [Oh3] and [OS1] (see also [GS]). Throughout the article, let M be a connected, n -dimensional \mathcal{C}^∞ -manifold without boundary such that $n \geq 2$. We fix an arbitrary positive \mathcal{C}^∞ -measure m on M as our base measure.

2.1 Finsler manifolds

Given a local coordinate $(x^i)_{i=1}^n$ on an open set $\Omega \subset M$, we will always use the coordinate $(x^i, v^j)_{i,j=1}^n$ of $T\Omega$ such that

$$v = \sum_{j=1}^n v^j \frac{\partial}{\partial x^j} \Big|_x \in T_x M \quad \text{for } x \in \Omega.$$

Definition 2.1 (Finsler structures) A nonnegative function $F : TM \rightarrow [0, \infty)$ is called a \mathcal{C}^∞ -Finsler structure of M if the following three conditions hold.

- (1) (Regularity) F is \mathcal{C}^∞ on $TM \setminus 0$, where 0 stands for the zero section.
- (2) (Positive 1-homogeneity) It holds $F(cv) = cF(v)$ for all $v \in TM$ and $c > 0$.
- (3) (Strong convexity) The $n \times n$ matrix

$$(g_{ij}(v))_{i,j=1}^n := \left(\frac{1}{2} \frac{\partial^2 (F^2)}{\partial v^i \partial v^j} (v) \right)_{i,j=1}^n \quad (2.1)$$

is positive-definite for all $v \in TM \setminus 0$.

We call such a pair (M, F) a \mathcal{C}^∞ -Finsler manifold.

That is to say, $F|_{T_x M}$ is a smooth *Minkowski norm* for every $x \in M$, and F varies smoothly also in the horizontal direction. We will denote the *unit tangent sphere bundle* by $UM := TM \cap F^{-1}(1)$. For $x, y \in M$, we define the *distance* from x to y in a natural way by

$$d(x, y) := \inf_{\eta} \int_0^1 F(\dot{\eta}(t)) dt,$$

where the infimum is taken over all \mathcal{C}^1 -curves $\eta : [0, 1] \rightarrow M$ such that $\eta(0) = x$ and $\eta(1) = y$. We remark that our distance can be *nonsymmetric* (namely $d(y, x) \neq d(x, y)$) since F is only positively homogeneous. A \mathcal{C}^∞ -curve η on M is called a *geodesic* if it is locally minimizing and has a constant speed (i.e., $F(\dot{\eta})$ is constant). See (2.7) below for the precise geodesic equation. Given $v \in T_x M$, if there is a geodesic $\eta : [0, 1] \rightarrow M$ with $\dot{\eta}(0) = v$, then we define the *exponential map* by $\exp_x(v) := \eta(1)$. We say that (M, F) is *forward complete* if the exponential map is defined on whole TM . Then by the Hopf-Rinow theorem any pair of points is connected by a minimal geodesic (cf. [BCS, Theorem 6.6.1]).

For each $v \in T_x M \setminus 0$, the positive-definite matrix $(g_{ij}(v))_{i,j=1}^n$ in (2.1) induces the Riemannian structure g_v of $T_x M$ as

$$g_v \left(\sum_{i=1}^n a_i \frac{\partial}{\partial x^i} \Big|_x, \sum_{j=1}^n b_j \frac{\partial}{\partial x^j} \Big|_x \right) := \sum_{i,j=1}^n a_i b_j g_{ij}(v). \quad (2.2)$$

This inner product is regarded as the best Riemannian approximation of $F|_{T_x M}$ in the direction v , and plays a vital role in the Riemannian geometric approach to Finsler geometry. A geometric way of introducing g_v is that the unit sphere of g_v is tangent to that of $F|_{T_x M}$ at $v/F(v)$ up to the second order. In particular, we have $g_v(v, v) = F(v)^2$.

For later convenience, we recall a useful fact on homogeneous functions.

Theorem 2.2 (cf. [BCS, Theorem 1.2.1]) *Suppose that a differentiable function $H : \mathbb{R}^n \setminus \{0\} \rightarrow \mathbb{R}$ satisfies $H(cv) = c^r H(v)$ for some $r \in \mathbb{R}$ and all $c > 0$ and $v \in \mathbb{R}^n \setminus \{0\}$ (that is, H is positively r -homogeneous). Then we have*

$$\sum_{i=1}^n \frac{\partial H}{\partial v^i}(v) v^i = rH(v) \quad \text{for all } v \in \mathbb{R}^n \setminus \{0\}.$$

The *Cartan tensor*

$$A_{ijk}(v) := \frac{F(v)}{2} \frac{\partial g_{ij}}{\partial v^k}(v) \quad \text{for } v \in TM \setminus 0$$

is a quantity appearing only in the Finsler context. Indeed, A_{ijk} vanishes everywhere on $TM \setminus 0$ if and only if F comes from a Riemannian metric. As g_{ij} is positively 0-homogeneous on each $T_x M \setminus 0$, Theorem 2.2 yields

$$\sum_{i=1}^n A_{ijk}(v) v^i = \sum_{j=1}^n A_{ijk}(v) v^j = \sum_{k=1}^n A_{ijk}(v) v^k = 0 \quad (2.3)$$

for all $v \in TM \setminus 0$ and $i, j, k = 1, 2, \dots, n$. Define the *formal Christoffel symbol*

$$\gamma_{jk}^i(v) := \frac{1}{2} \sum_{l=1}^n g^{il}(v) \left\{ \frac{\partial g_{jl}}{\partial x^k}(v) + \frac{\partial g_{lk}}{\partial x^j}(v) - \frac{\partial g_{jk}}{\partial x^l}(v) \right\} \quad \text{for } v \in TM \setminus 0,$$

where $(g^{ij}(v))$ stands for the inverse matrix of $(g_{ij}(v))$. We also introduce the *geodesic spray coefficient* and the *nonlinear connection*

$$G^i(v) := \sum_{j,k=1}^n \gamma_{jk}^i(v) v^j v^k, \quad N_j^i(v) := \frac{1}{2} \frac{\partial G^i}{\partial v^j}(v) \quad \text{for } v \in TM \setminus 0,$$

and $G^i(0) = N_j^i(0) := 0$ by convention. Note that G^i is positively 2-homogeneous, so that Theorem 2.2 implies $\sum_{j=1}^n N_j^i(v) v^j = G^i(v)$. Following another representation of N_j^i (with the help of Theorem 2.2) would be useful,

$$N_j^i(v) = \sum_{k=1}^n \gamma_{jk}^i(v) v^k - \frac{1}{F(v)} \sum_{k,l,m=1}^n A_{jk}^i(v) \gamma_{lm}^k(v) v^l v^m, \quad (2.4)$$

where $A_{jk}^i := \sum_{l=1}^n g^{il} A_{ljk}$.

By using N_j^i , the coefficients of the *Chern connection* are given by

$$\Gamma_{jk}^i := \gamma_{jk}^i - \sum_{l,m=1}^n \frac{g^{il}}{F} (A_{jlm} N_k^m + A_{lkm} N_j^m - A_{jkm} N_l^m) \quad \text{on } TM \setminus 0. \quad (2.5)$$

That is, the corresponding *covariant derivative* of a vector field $X = \sum_{i=1}^n X^i(\partial/\partial x^i)$ by $v \in T_x M$ with *reference vector* $w \in T_x M \setminus 0$ is defined as

$$D_v^w X(x) := \sum_{i,j=1}^n \left\{ v^j \frac{\partial X^i}{\partial x^j}(x) + \sum_{k=1}^n \Gamma_{jk}^i(w) v^j X^k(x) \right\} \frac{\partial}{\partial x^i} \Big|_x \in T_x M. \quad (2.6)$$

Then the *geodesic equation* is written as, with the help of (2.3),

$$D_{\dot{\eta}}^{\dot{\eta}} \dot{\eta}(t) = \sum_{i=1}^n \{ \ddot{\eta}^i(t) + G^i(\dot{\eta}(t)) \} \frac{\partial}{\partial x^i} \Big|_{\eta(t)} = 0. \quad (2.7)$$

The following fact will be used in Section 5, we give a proof for completeness.

Lemma 2.3 *If all integral curves of a non-vanishing C^∞ -vector field V are geodesic, then we have*

$$D_V^V W = D_V^{g_V} W, \quad D_W^V V = D_W^{g_V} V$$

for any differentiable vector field W , where D^{g_V} stands for the covariant derivative with respect to the Riemannian structure g_V given as (2.2).

Proof. To see the claim, it suffices to compare $\sum_{j,k=1}^n \Gamma_{jk}^i(V) W^j V^k$ with that for g_V . On the one hand, we observe from (2.5) and (2.3) that

$$\sum_{j,k=1}^n \Gamma_{jk}^i(V) W^j V^k = \sum_{j=1}^n \left\{ \sum_{k=1}^n \gamma_{jk}^i(V) V^k - \sum_{l,m=1}^n \frac{g^{il}(V)}{F(V)} A_{jlm}(V) G^m(V) \right\} W^j.$$

On the other hand, since

$$\frac{\partial [g_{jl}(V)]}{\partial x^k} = \frac{\partial g_{jl}(V)}{\partial x^k} + \sum_{m=1}^n \frac{2}{F(V)} A_{jlm}(V) \frac{\partial V^m}{\partial x^k},$$

the corresponding quantity for g_V is

$$\sum_{j,k=1}^n \left\{ \gamma_{jk}^i(V) + \sum_{l,m=1}^n \frac{g^{il}(V)}{F(V)} A_{jlm}(V) \frac{\partial V^m}{\partial x^k} \right\} W^j V^k.$$

Then the geodesic equation

$$\sum_{k=1}^n \frac{\partial V^m}{\partial x^k} V^k + G^m(V) = 0$$

shows that they coincide. □

2.2 Weighted Ricci curvature

The *Ricci curvature* (as the trace of the *flag curvature*) for a Finsler manifold is defined by using the Chern connection. Instead of giving the precise definition in coordinates, we explain an elegant interpretation due to Shen ([Sh1, §6.2], [Sh2, Lemma 2.4]).

Given a unit vector $v \in U_x M$, we extend it to a C^∞ -vector field V on a neighborhood of x in such a way that every integral curve of V is geodesic, and consider the Riemannian structure g_V induced from (2.2). Then the flag curvature $\mathbf{K}(v, w)$ for $w \in T_x M$ linearly independent with v coincides with the sectional curvature of the plane spanned by v and w with respect to g_V (in particular, it is independent of the choice of V). Similarly, the

Ricci curvature $\text{Ric}(v)$ of v with respect to F coincides with the Ricci curvature of v with respect to g_V .

Inspired by the above interpretation of the Ricci curvature and the theory of weighted Riemannian manifolds, the weighted Ricci curvature for the triple (M, F, m) was introduced in [Oh3] as follows.

Definition 2.4 (Weighted Ricci curvature) We first define the function $\Psi : UM \rightarrow \mathbb{R}$ on the unit tangent sphere bundle via the decomposition $m = e^{-\Psi(\dot{\eta})} \text{vol}_{\dot{\eta}}$ along unit speed geodesics η , where $\text{vol}_{\dot{\eta}}$ denotes the Riemannian volume measure of $g_{\dot{\eta}}$. Then, given a unit vector $v \in U_x M$ and the geodesic $\eta : (-\varepsilon, \varepsilon) \rightarrow M$ such that $\dot{\eta}(0) = v$, we define the *weighted Ricci curvature* involving a parameter $N \in [n, \infty]$ by

$$(1) \quad \text{Ric}_n(v) := \begin{cases} \text{Ric}(v) + (\Psi \circ \dot{\eta})''(0) & \text{if } (\Psi \circ \dot{\eta})'(0) = 0, \\ -\infty & \text{if } (\Psi \circ \dot{\eta})'(0) \neq 0, \end{cases}$$

$$(2) \quad \text{Ric}_N(v) := \text{Ric}(v) + (\Psi \circ \dot{\eta})''(0) - \frac{(\Psi \circ \dot{\eta})'(0)^2}{N - n} \quad \text{for } N \in (n, \infty),$$

$$(3) \quad \text{Ric}_\infty(v) := \text{Ric}(v) + (\Psi \circ \dot{\eta})''(0).$$

We also set $\text{Ric}_N(cv) := c^2 \text{Ric}_N(v)$ for $c \geq 0$.

We will say that $\text{Ric}_N \geq K$ holds for some $K \in \mathbb{R}$ if $\text{Ric}_N(v) \geq KF(v)^2$ for all $v \in TM$. We remark that $(\Psi \circ \dot{\eta})'(0)$ coincides with Shen's \mathbf{S} -curvature $\mathbf{S}(v)$ (see [Sh1, §7.3]). Observe that $\text{Ric}_N(v) \leq \text{Ric}_{N'}(v)$ holds for $N < N'$. It was shown in [Oh3, Theorem 1.2] that, for each $K \in \mathbb{R}$, $\text{Ric}_N \geq K$ is equivalent to Lott, Sturm and Villani's *curvature-dimension condition* $\text{CD}(K, N)$. This equivalence extends the corresponding result on (weighted) Riemannian manifolds (due to [vRS], [St1], [St2], [St3], [LV1], [LV2]), and has many analytic and geometric applications (see [Oh3]).

Remark 2.5 For a Riemannian manifold (M, g, vol_g) endowed with the Riemannian volume measure, clearly we have $\Psi \equiv 0$ and hence $\text{Ric}_N = \text{Ric}$ for all $N \in [n, \infty]$. In general, however, a Finsler manifold may not admit any measure m satisfying $\mathbf{S} \equiv 0$ (in other words, $\text{Ric}_n > -\infty$), see [Oh4] for such an example. This means that there is no nice reference measure, so that we began with an arbitrary measure.

For later convenience, we introduce the following notations.

Definition 2.6 (Reverse Finsler structure) Define the *reverse Finsler structure* \overleftarrow{F} of F by $\overleftarrow{F}(v) := F(-v)$. We will put an arrow \leftarrow on those quantities associated with \overleftarrow{F} , for example, $\overleftarrow{d}(x, y) = d(y, x)$, $\overleftarrow{\nabla}u = -\nabla(-u)$ and $\overleftarrow{\text{Ric}}_N(v) = \text{Ric}_N(-v)$.

Note that $\text{Ric}_N \geq K$ is equivalent to $\overleftarrow{\text{Ric}}_N(v) \geq K\overleftarrow{F}(v)^2$, so that the Ricci curvature bound is equivalent between F and \overleftarrow{F} . We say that (M, F) is *backward complete* if (M, \overleftarrow{F}) is forward complete. The forward and backward completeness are not mutually equivalent in general.

2.3 Nonlinear Laplacian

Let us denote by $\mathcal{L}^* : T^*M \rightarrow TM$ the *Legendre transform* associated with F and its dual norm F^* on T^*M . Precisely, \mathcal{L}^* is sending $\alpha \in T_x^*M$ to the unique element $v \in T_xM$ such that $\alpha(v) = F^*(\alpha)^2$ and $F(v) = F^*(\alpha)$. Note that $\mathcal{L}^*|_{T_x^*M}$ becomes a linear operator only when $F|_{T_xM}$ is an inner product. For a differentiable function $u : M \rightarrow \mathbb{R}$, the *gradient vector* of u at x is defined as the Legendre transform of the derivative, $\nabla u(x) := \mathcal{L}^*(Du(x)) \in T_xM$. For a differentiable vector field V on M and $x \in M$ such that $V(x) \neq 0$, we define $\nabla V(x) \in T_x^*M \otimes T_xM$ by using the covariant derivative (2.6) as

$$\nabla V(v) := D_v^V V \in T_xM \quad \text{for } v \in T_xM.$$

We also set $\nabla^2 u(x) := \nabla(\nabla u)(x)$ for a twice differentiable function $u : M \rightarrow \mathbb{R}$ and $x \in M$ such that $Du(x) \neq 0$.

Define the *divergence* of a differentiable vector field $V = \sum_{i=1}^n V^i(\partial/\partial x^i)$ on M with respect to the base measure m by

$$\operatorname{div}_m V := \sum_{i=1}^n \left(\frac{\partial V^i}{\partial x^i} + V^i \frac{\partial \Phi}{\partial x^i} \right),$$

where we decomposed m in coordinates as $dm = e^\Phi dx^1 dx^2 \cdots dx^n$. The divergence can be rewritten (and extended to weakly differentiable vector fields) in the weak form as

$$\int_M \phi \operatorname{div}_m V \, dm = - \int_M D\phi(V) \, dm \quad \text{for all } \phi \in \mathcal{C}_c^\infty(M).$$

Then we define the distributional *Laplacian* of $u \in H_{\text{loc}}^1(M)$ by $\Delta u := \operatorname{div}_m(\nabla u)$ in the weak sense that

$$\int_M \phi \Delta u \, dm := - \int_M D\phi(\nabla u) \, dm \quad \text{for all } \phi \in \mathcal{C}_c^\infty(M).$$

We remark that $H_{\text{loc}}^1(M)$ is defined solely in terms of the differentiable structure of M . As the Legendre transform is nonlinear, this Laplacian is a nonlinear operator unless F comes from a Riemannian metric.

The weighted Ricci curvature Ric_N works quite well with the nonlinear Laplacian. Among others, we recall the Laplacian comparison theorem ([OS1, Theorem 5.2]) in the special case of nonnegative curvature, as well as the Bochner-Weitzenböck formula ([OS3, Theorems 3.3, 3.6]).

Theorem 2.7 (Laplacian comparison theorem) *Let (M, F) be forward or backward complete, and assume that $\operatorname{Ric}_N \geq 0$ for some $N \in [n, \infty)$. Then, for any $z \in M$, the function $u(x) = d(z, x)$ satisfies*

$$\Delta u(x) \leq \frac{N-1}{d(z, x)}$$

point-wise on $M \setminus (\{z\} \cup \operatorname{Cut}_z)$, and in the distributional sense on $M \setminus \{z\}$.

We denote by Cut_z the *cut locus* of z . The cut locus is the set of *cut points* $x = \exp_z(v)$ such that $\eta(t) = \exp_z(tv)$ is minimal on $[0, 1]$ but not minimal on $[0, 1 + \varepsilon]$ for any $\varepsilon > 0$.

Theorem 2.8 (Bochner-Weitzenböck formula) *Given $u \in H_{\text{loc}}^2(M) \cap \mathcal{C}^1(M)$ with $\Delta u \in H_{\text{loc}}^1(M)$, we have*

$$\Delta^{\nabla u} \left(\frac{F(\nabla u)^2}{2} \right) - D(\Delta u)(\nabla u) = \text{Ric}_{\infty}(\nabla u) + \|\nabla^2 u\|_{HS(\nabla u)}^2$$

as well as

$$\Delta^{\nabla u} \left(\frac{F(\nabla u)^2}{2} \right) - D(\Delta u)(\nabla u) \geq \text{Ric}_N(\nabla u) + \frac{(\Delta u)^2}{N}$$

for $N \in [n, \infty]$ point-wise on $M \setminus \{x \in M \mid \nabla u(x) = 0\}$, and in the weak sense on M . Here $\Delta^{\nabla u} := \text{div}_m \circ \nabla^{g_{\nabla u}}$ is the linearized Laplacian associated with $g_{\nabla u}$, and $\|\cdot\|_{HS(\nabla u)}$ stands for the Hilbert-Schmidt norm with respect to $g_{\nabla u}$.

To be precise, in the definition of $\Delta^{\nabla u}$, we replace ∇u with a measurable, non-vanishing vector field V such that $V(x) = \nabla u(x)$ if $\nabla u(x) \neq 0$. We remark that $\Delta^{\nabla u} u = \Delta u$ holds ([OS1, Lemma 2.4]).

2.4 Berwald spaces

We introduce an important and reasonable class of Finsler manifolds.

Definition 2.9 (Berwald spaces) We say that a Finsler manifold (M, F) is of *Berwald type* (or a *Berwald space*) if Γ_{jk}^i is constant on $T_x M \setminus 0$ for every $x \in M$.

Clearly Riemannian manifolds and (smooth) Minkowski normed spaces are of Berwald type. Berwald spaces enjoy several fine properties (see [BCS, Chapter 10]), we recall two of them for later use (cf. [BCS, Proposition 10.1.1, Theorem 10.6.2]).

Proposition 2.10 (Isometry of tangent spaces, [Ic]) *Let (M, F) be a Finsler manifold of Berwald type. Then, for any \mathcal{C}^1 -curve $\eta : [0, 1] \rightarrow M$ with $\dot{\eta} \neq 0$, the parallel transport along η is a linear isometry between $(T_{\eta(0)}M, F|_{T_{\eta(0)}M})$ and $(T_{\eta(1)}M, F|_{T_{\eta(1)}M})$.*

We remark that, in Berwald spaces, the covariant derivative (2.6) is independent of the choice of a reference vector. Thus the parallel transport is unambiguously defined.

Theorem 2.11 (Szabó's rigidity, [Sz1]) *Let (M, F) be a connected Berwald surface. Then the following dichotomy holds.*

- *If the flag curvature is identically 0, then F is locally Minkowskian everywhere.*
- *If the flag curvature is not identically 0, then F is Riemannian everywhere.*

Szabó also classified higher dimensional non-Riemannian Berwald metrics by means of holonomy theory ([Sz1], [Sz2]), whereas such a classification is not really helpful to our purpose. We also remark that the *Busemann-Hausdorff measure* satisfies $\mathbf{S} \equiv 0$ for Berwald spaces (cf. [Sh1, §7.3]), though this fact will not be used.

3 Analysis of Busemann functions

Let (M, F) be forward complete in this section. We begin the study of the splitting phenomenon with analyzing Busemann functions.

We call a geodesic $\eta : [0, \infty) \rightarrow M$ a *ray* if it is globally minimizing and has the unit speed (for brevity). Given a ray η , the associated *Busemann function* $\mathbf{b}_\eta : M \rightarrow \mathbb{R}$ is defined by

$$\mathbf{b}_\eta(x) := \lim_{t \rightarrow \infty} \{t - d(x, \eta(t))\}.$$

This limit indeed exists because the triangle inequality ensures, for any $s < t$,

$$s - d(x, \eta(s)) \leq s - \{d(x, \eta(t)) - (t - s)\} = t - d(x, \eta(t)) \leq d(\eta(0), x).$$

The triangle inequality also shows that \mathbf{b}_η is 1-Lipschitz in the sense that

$$\mathbf{b}_\eta(y) - \mathbf{b}_\eta(x) \leq d(x, y) \quad \text{for all } x, y \in M, \quad (3.1)$$

and hence \mathbf{b}_η is differentiable almost everywhere.

We say that another ray $\sigma : [0, \infty) \rightarrow M$ is *asymptotic* to η , denoted by $\sigma \sim \eta$, if there are sequences $\{t_i\}_{i \in \mathbb{N}} \subset [0, \infty)$ and $\{\sigma_i\}_{i \in \mathbb{N}}$ such that $\lim_{i \rightarrow \infty} t_i = \infty$, $\sigma_i : [0, d(\sigma(0), \eta(t_i))] \rightarrow M$ is a minimal geodesic from $\sigma(0)$ to $\eta(t_i)$, and that $\lim_{i \rightarrow \infty} \sigma_i(t) = \sigma(t)$ for all $t \geq 0$. The next lemma is concerned with the fundamental properties of Busemann functions (cf. [SST, Theorem 3.8.2]). We give proofs for completeness as our distance is nonsymmetric.

Lemma 3.1 *Let $\eta : [0, \infty) \rightarrow M$ be a ray.*

- (i) *For any $x \in M$, there exists a ray σ asymptotic to η such that $\sigma(0) = x$.*
- (ii) *For any ray $\sigma \sim \eta$ and $s \geq 0$, it holds $\mathbf{b}_\eta(\sigma(s)) = \mathbf{b}_\eta(\sigma(0)) + s$.*
- (iii) *If \mathbf{b}_η is differentiable at $x \in M$, then $\sigma(s) = \exp_x(s \nabla \mathbf{b}_\eta(x))$ is a unique ray asymptotic to η emanating from x .*

Proof. (i) As (M, F) is forward complete, we can choose a unit speed minimal geodesic σ_i from x to $\eta(i)$ for each $i \in \mathbb{N}$. By extracting a subsequence denoted again by $\{\sigma_i\}_{i \in \mathbb{N}}$, the initial tangent vector $\dot{\sigma}_i(0)$ converges to some unit vector $v \in U_x M$. Then the ray $\sigma(s) := \exp_x(sv)$ is asymptotic to η by construction.

(ii) Take $\{t_i\}_{i \in \mathbb{N}}$ and $\{\sigma_i\}_{i \in \mathbb{N}}$ as in the definition of the asymptoticity. Then it holds

$$\mathbf{b}_\eta(\sigma(s)) = \lim_{i \rightarrow \infty} \{t_i - d(\sigma(s), \eta(t_i))\}$$

by the definition of \mathbf{b}_η . We can replace $\sigma(s)$ in the right hand side with $\sigma_i(s)$ since

$$|d(\sigma(s), \eta(t_i)) - d(\sigma_i(s), \eta(t_i))| \leq \max \{d(\sigma(s), \sigma_i(s)), d(\sigma_i(s), \sigma(s))\} \rightarrow 0$$

as $i \rightarrow \infty$. Hence we have, by the choice of σ_i ,

$$\begin{aligned} \mathbf{b}_\eta(\sigma(s)) &= \lim_{i \rightarrow \infty} \{t_i - d(\sigma_i(s), \eta(t_i))\} = \lim_{i \rightarrow \infty} \{t_i - d(\sigma_i(0), \eta(t_i)) + s\} \\ &= \lim_{i \rightarrow \infty} \{t_i - d(\sigma(0), \eta(t_i)) + s\} = \mathbf{b}_\eta(\sigma(0)) + s. \end{aligned}$$

(iii) Recall from (3.1) that \mathbf{b}_η is 1-Lipschitz. Then we deduce from (ii) that any ray $\sigma \sim \eta$ with $\sigma(0) = x$ must satisfy $\dot{\sigma}(0) = \nabla \mathbf{b}_\eta(x)$. This completes the proof. \square

The following is a key analytic property of Busemann functions. The proof is similar to [KS, Lemma 5.6] (see also [EH], [FLZ, Lemma 2.1]) thanks to the Laplacian comparison theorem (Theorem 2.7).

Proposition 3.2 *Assume that $\text{Ric}_N \geq 0$ for some $N \in [n, \infty]$, and that $\Psi : UM \rightarrow \mathbb{R}$ as in Definition 2.4 is bounded above if $N = \infty$. Then \mathbf{b}_η is subharmonic, namely $\Delta \mathbf{b}_\eta \geq 0$ holds in the distributional sense.*

Proof. We first treat the case of $N < \infty$. Fix an arbitrary bounded open set $\Omega \subset M$ and a nonnegative test function $\phi \in H_0^1(\Omega)$. Put $r_i(x) := -d(x, \eta(i))$ for $i \in \mathbb{N}$. Note that r_i is differentiable almost everywhere and $\nabla r_i(x)$ coincides with the initial vector of the unique unit speed minimal geodesic from x to $\eta(i)$. Thanks to Lemma 3.1(iii) (and the construction in (i)), we find $\lim_{i \rightarrow \infty} \nabla r_i(x) = \nabla \mathbf{b}_\eta(x)$ for x at where \mathbf{b}_η is differentiable. Thus we have, by the dominated convergence theorem,

$$\lim_{i \rightarrow \infty} \int_{\Omega} D\phi(\nabla r_i) dm = \int_{\Omega} D\phi(\nabla \mathbf{b}_\eta) dm.$$

In order to apply Theorem 2.7, we observe (recall Definition 2.6)

$$\nabla r_i = -\overleftarrow{\nabla}(-r_i) = -\overleftarrow{\nabla} \left[\overleftarrow{d}(\eta(i), \cdot) \right].$$

Hence Theorem 2.7 for \overleftarrow{F} yields

$$\int_{\Omega} D\phi(\nabla \mathbf{b}_\eta) dm = \lim_{i \rightarrow \infty} \int_{\Omega} \phi \overleftarrow{\Delta}(-r_i) dm \leq (N-1) \lim_{i \rightarrow \infty} \int_{\Omega} \frac{\phi}{-r_i} dm = 0. \quad (3.2)$$

As for $N = \infty$, we derive from the calculation with respect to the Riemannian structure $g_{\nabla r_i}$ that (cf. [FLZ, (2.1)]), since $\overleftarrow{g}_{\overleftarrow{\nabla}(-r_i)} = g_{-\overleftarrow{\nabla}(-r_i)} = g_{\nabla r_i}$ and all integral curves of ∇r_i are geodesic (with respect to F),

$$\begin{aligned} \overleftarrow{\Delta}(-r_i)(x) &\leq -\frac{n-1}{r_i(x)} + \frac{2\Psi(\dot{\sigma}(0))}{r_i(x)} + \frac{2}{r_i(x)^2} \int_0^{-r_i(x)} \Psi(\dot{\sigma}) ds \\ &\leq -\frac{1}{r_i(x)} \left\{ (n-1) - 2\Psi(\dot{\sigma}(0)) + 2 \sup_{UM} \Psi \right\} \end{aligned}$$

for $x \in M \setminus (\{\eta(i)\} \cup \overleftarrow{\text{Cut}}_{\eta(i)})$, where $\sigma : [0, -r_i(x)] \rightarrow M$ is the unique minimal geodesic from x to $\eta(i)$ (with respect to F). Therefore (3.2) is available with

$$N_{\Omega} = n + 2 \left(\sup_{UM} \Psi - \inf_{U\Omega} \Psi \right)$$

in place of N , and \mathbf{b}_η is subharmonic. \square

4 A diffeomorphic splitting

From here on, let (M, F) be both forward and backward complete, and assume that $\text{Ric}_N \geq 0$ for some $N \in [n, \infty]$ and Ψ (defined in Definition 2.4) is bounded above if $N = \infty$. Suppose that (M, F) admits a *straight line*, that is, a unit speed, globally minimizing geodesic $\eta : \mathbb{R} \rightarrow M$. Let us consider the two Busemann functions

$$\mathbf{b}_\eta(x) := \lim_{t \rightarrow \infty} \{t - d(x, \eta(t))\}, \quad \mathbf{b}_{\bar{\eta}}(x) := \lim_{t \rightarrow \infty} \{t - d(\eta(-t), x)\},$$

where $\mathbf{b}_{\bar{\eta}}$ is precisely the Busemann function for the ray $\bar{\eta}(t) := \eta(-t)$, $t \in [0, \infty)$, with respect to \overleftarrow{F} .

Proposition 4.1 *Let $\eta : \mathbb{R} \rightarrow M$ be a straight line. Then we have $\mathbf{b}_\eta + \mathbf{b}_{\bar{\eta}} \equiv 0$, and \mathbf{b}_η and $\mathbf{b}_{\bar{\eta}}$ are harmonic with respect to F and \overleftarrow{F} , namely $\Delta \mathbf{b}_\eta = \overleftarrow{\Delta} \mathbf{b}_{\bar{\eta}} \equiv 0$. In particular, \mathbf{b}_η and $\mathbf{b}_{\bar{\eta}}$ are \mathcal{C}^∞ and $\Delta \mathbf{b}_\eta = \overleftarrow{\Delta} \mathbf{b}_{\bar{\eta}} \equiv 0$ in fact holds in the point-wise sense.*

Proof. We immediately observe from the triangle inequality that $\mathbf{b}_\eta + \mathbf{b}_{\bar{\eta}} \leq 0$. Proposition 3.2 implies $\Delta \mathbf{b}_\eta \geq 0$ as well as $\overleftarrow{\Delta} \mathbf{b}_{\bar{\eta}} \geq 0$ (note that $\overleftarrow{\Psi}(v) = \Psi(-v)$). Therefore

$$\Delta \mathbf{b}_\eta \geq 0 \geq -\overleftarrow{\Delta} \mathbf{b}_{\bar{\eta}} = \Delta(-\mathbf{b}_{\bar{\eta}}),$$

while $\mathbf{b}_\eta \circ \eta \equiv -\mathbf{b}_{\bar{\eta}} \circ \eta$. Hence the strong maximum principle (see [Da, Theorem 2-2], [GS, Lemma 5.4]) yields $\mathbf{b}_\eta = -\mathbf{b}_{\bar{\eta}}$ and $\Delta \mathbf{b}_\eta = \overleftarrow{\Delta} \mathbf{b}_{\bar{\eta}} \equiv 0$.

As a harmonic function is a static solution to the heat equation, \mathbf{b}_η is $\mathcal{C}^{1,\alpha}$ by [OS1, Theorem 4.9] (see also [GS, Theorem 1.1]). Furthermore, $\nabla \mathbf{b}_\eta$ does not vanish since $F(\nabla \mathbf{b}_\eta) \equiv 1$ by Lemma 3.1(ii), so that \mathbf{b}_η and $\mathbf{b}_{\bar{\eta}} = -\mathbf{b}_\eta$ are eventually \mathcal{C}^∞ (see [OS1, Remark 4.10], [GS, Theorem 1.1]). \square

We say that a straight line $\sigma : \mathbb{R} \rightarrow M$ is *bi-asymptotic* to η if $\sigma|_{[0,\infty)} \sim \eta|_{[0,\infty)}$ and if $\bar{\sigma}(s) := \sigma(-s)$ is asymptotic to $\bar{\eta}$ with respect to \overleftarrow{F} . Combining Proposition 4.1 with Lemma 3.1(iii), we observe the following.

Lemma 4.2 *Let $\eta : \mathbb{R} \rightarrow M$ be a straight line. Then, for any $x \in M$, the geodesic $\sigma : \mathbb{R} \rightarrow M$ with $\dot{\sigma}(0) = \nabla \mathbf{b}_\eta(x)$ is a unique straight line bi-asymptotic to η such that $\sigma(0) = x$.*

Lemma 3.1(ii) implies not only $\nabla \mathbf{b}_\eta \neq 0$ but also that every integral curve of $\nabla \mathbf{b}_\eta$ is geodesic. Therefore $\text{Ric}_N(\nabla \mathbf{b}_\eta) = \text{Ric}_N^{g_{\nabla \mathbf{b}_\eta}}(\nabla \mathbf{b}_\eta)$ and we can apply the Cheeger-Gromoll-Lichnerowicz splitting theorem ([CG1], [Li]) to the weighted Riemannian manifold $(M, g_{\nabla \mathbf{b}_\eta}, m)$.

Proposition 4.3 (Isometric splitting of $(M, g_{\nabla \mathbf{b}_\eta}, m)$) *If (M, F) contains a straight line $\eta : \mathbb{R} \rightarrow M$, then $(M, g_{\nabla \mathbf{b}_\eta})$ splits isometrically as $M = M' \times \mathbb{R}$ with $M' = \mathbf{b}_\eta^{-1}(0)$, and $\Psi \circ \dot{\sigma}$ is constant on the line $\sigma(s) = (x, s) \in M' \times \mathbb{R}$ for each $x \in M'$.*

Proof. We give a sketch of the proof for thoroughness, see [FLZ, Theorem 1.1], [WW, Theorem 6.1] for details. Applying the Bochner-Weitzenböck formula for $g_{\nabla \mathbf{b}_\eta}$ to \mathbf{b}_η , we deduce from $\Delta^{\nabla \mathbf{b}_\eta} \mathbf{b}_\eta = \Delta \mathbf{b}_\eta \equiv 0$ that

$$\mathrm{Ric}_\infty(\nabla \mathbf{b}_\eta) + \|\nabla^2 \mathbf{b}_\eta\|_{HS}^2 = \Delta \left(\frac{|\nabla \mathbf{b}_\eta|^2}{2} \right) - D(\Delta \mathbf{b}_\eta)(\nabla \mathbf{b}_\eta) = 0, \quad (4.1)$$

where ∇ and Δ are with respect to $g_{\nabla \mathbf{b}_\eta}$ and m . Thus the hypothesis $\mathrm{Ric}_\infty \geq \mathrm{Ric}_N \geq 0$ shows $\nabla^2 \mathbf{b}_\eta \equiv 0$, namely $\nabla \mathbf{b}_\eta$ is a parallel (and hence Killing) vector field. Therefore the associated one-parameter family of transforms $\varphi_t : M \rightarrow M$, $t \in \mathbb{R}$, consists of isometries (with respect to $g_{\nabla \mathbf{b}_\eta}$) and M is isometric to $M' \times \mathbb{R}$ with $M' := \mathbf{b}_\eta^{-1}(0)$.

In order to split the measure m , we observe from

$$\|\nabla^2 \mathbf{b}_\eta\|_{HS}^2 \geq \frac{(\Delta \mathbf{b}_\eta + D(\Psi \circ \nabla \mathbf{b}_\eta)(\nabla \mathbf{b}_\eta))^2}{n}$$

that $D(\Psi \circ \nabla \mathbf{b}_\eta)(\nabla \mathbf{b}_\eta) \equiv 0$ (see, for example, the calculation in the proof of [OS3, Theorem 3.3]). \square

The boundedness of Ψ is in fact necessary for splitting both $g_{\nabla \mathbf{b}_\eta}$ and m , see [WW, Examples 2.1, 2.2] for simple interesting examples. Recall that we used the boundedness only in Proposition 3.2.

Corollary 4.4 (Diffeomorphic splitting of (M, m)) *Assume that (M, F) includes a straight line. Then (M, m) admits a diffeomorphic, measure-preserving splitting $(M, m) = (M' \times \mathbb{R}, m' \times \mathbf{L}^1)$, where \mathbf{L}^1 is the one-dimensional Lebesgue measure and $m' := m|_{M'}$.*

To be precise, the map

$$(M' \times \mathbb{R}, m' \times \mathbf{L}^1) \ni (x, t) \mapsto \varphi_t(x) \in (M, m)$$

is diffeomorphic and measure-preserving. It is unclear if this procedure can be iterated, because it seems difficult to determine the structures of $(M', F|_{TM'}, m')$ as well as $(M', g_{\nabla \mathbf{b}_\eta}|_{TM'}, m')$ from the construction in Proposition 4.3 (for instance, we do not know if they have the nonnegative curvature).

Remark 4.5 Corollary 4.4 would be compared with the structure theorem for a (non-branching) metric measure space (X, d, m) satisfying the *measure contraction property* $\mathrm{MCP}(K, N)$ (see [Oh1], [St3, Section 5]) with $N \in (1, \infty)$ and $K > 0$, as $\mathrm{CD}(K, N)$ implies $\mathrm{MCP}(K, N)$. If (X, d) attains the maximal diameter $\pi\sqrt{(N-1)/K}$, then (X, m) is represented as the spherical suspension of some topological measure space ([Oh2, Section 5]). It is unknown whether a splitting theorem similar to Corollary 4.4 holds for general metric measure spaces satisfying $\mathrm{MCP}(0, N)$ or $\mathrm{CD}(0, N)$.

Though the following theorem is essentially included in Proposition 4.3, we state it separately for future convenience.

Theorem 4.6 *Suppose that (M, F, m) is forward and backward complete and satisfies $\mathrm{Ric}_N \geq 0$ for some $N \in [n, \infty]$, and that Ψ is bounded above if $N = \infty$. If (M, F) contains a straight line $\eta : \mathbb{R} \rightarrow M$, then we have $\nabla^2 \mathbf{b}_\eta \equiv 0$.*

Proof. Applying the Bochner-Weitzenböck formula (Theorem 2.8) to the harmonic function \mathbf{b}_η , we obtain

$$\text{Ric}_\infty(\nabla \mathbf{b}_\eta) + \|\nabla^2 \mathbf{b}_\eta\|_{HS(\nabla \mathbf{b}_\eta)}^2 = 0$$

and hence $\text{Ric}_\infty \geq \text{Ric}_N \geq 0$ implies $\nabla^2 \mathbf{b}_\eta \equiv 0$. (We remark that this formula in fact coincides with the Bochner-Weitzenböck formula (4.1) for $g_{\nabla \mathbf{b}_\eta}$ due to the fact that all integral curves of $\nabla \mathbf{b}_\eta$ are geodesic, see [OS3, Remark 3.4]. Thus we do not really need the formula in [OS3], the formula in the Riemannian case is enough.) \square

One may be able to derive from $\nabla^2 \mathbf{b}_\eta \equiv 0$ some more information on the structure of (M, F) , whereas we have succeeded only in the Berwald case (see the next section).

5 Berwald case

In this final section, we investigate a more detailed splitting phenomenon of Berwald spaces (recall Definition 2.9). Throughout the section, let (M, F, m) be a forward and backward complete Berwald space, and assume that $\text{Ric}_N \geq 0$ for some $N \in [n, \infty]$ and Ψ is bounded above if $N = \infty$. By the definition of Berwald spaces, the covariant derivative (2.6) does not depend on the choice of a reference vector, so that we will omit reference vectors in this section. In particular, the covariant derivative is linear in the sense that $D_v(W + X) = D_vW + D_vX$.

A subset $A \subset M$ is said to be *totally convex* if any minimal geodesic joining two points in A is contained in A . We say that $A \subset M$ is *geodesically complete* if, for any geodesic $\eta : (0, \varepsilon) \rightarrow M$ contained in A , its extension $\eta : \mathbb{R} \rightarrow M$ as a geodesic is still contained in A .

Lemma 5.1 *Suppose that (M, F) contains a straight line $\eta : \mathbb{R} \rightarrow M$. Then, given any geodesic $\xi : [0, l] \rightarrow M$, we have $(\mathbf{b}_\eta \circ \xi)'' \equiv 0$. In particular, for each $t \in \mathbb{R}$, $\mathbf{b}_\eta^{-1}(t)$ is totally convex and geodesically complete.*

Proof. We observe

$$(\mathbf{b}_\eta \circ \xi)'' = g_{\nabla \mathbf{b}_\eta}(D_{\dot{\xi}}(\nabla \mathbf{b}_\eta), \dot{\xi}) + g_{\nabla \mathbf{b}_\eta}(\nabla \mathbf{b}_\eta, D_{\dot{\xi}} \dot{\xi}) = 0$$

(see [BCS, Exercises 10.1.1, 10.1.2] for the first equality). To be precise, the first term vanishes in general by Theorem 4.6, while the second term vanishes only in Berwald spaces (since the covariant derivatives have $\nabla \mathbf{b}_\eta$ as the reference vector). We in particular find that $\mathbf{b}_\eta \circ \xi$ is constant if $\mathbf{b}_\eta(\xi(r)) = \mathbf{b}_\eta(\xi(s))$ for some $r \neq s$, so that $\mathbf{b}_\eta^{-1}(t)$ is totally convex and geodesically complete. \square

Define $\varphi_t : M \rightarrow M$, $t \in \mathbb{R}$, as the one-parameter family of \mathcal{C}^∞ -transforms generated from $\nabla \mathbf{b}_\eta$. Precisely, $\partial \varphi_t / \partial t = \nabla \mathbf{b}_\eta(\varphi_t)$.

Proposition 5.2 *Suppose that (M, F) contains a straight line $\eta : \mathbb{R} \rightarrow M$. Then we have the following.*

- (i) *For any $t \in \mathbb{R}$, φ_t is a measure-preserving isometry such that $\varphi_t(M_0) = M_t$, where we set $M_t := \mathbf{b}_\eta^{-1}(t)$. Moreover, it holds $M = \bigsqcup_{t \in \mathbb{R}} \varphi_t(M_0)$.*

- (ii) The $(n-1)$ -dimensional submanifold $(M_0, F|_{TM_0}, m_0)$ is again of Berwald type and satisfies $\text{Ric}_{N-1} \geq 0$, where $m_0 := m|_{M_0}$.
- (iii) Define the projection $\rho : M \rightarrow M_0$ by $\rho(\varphi_t(x)) := x$ for $(x, t) \in M_0 \times \mathbb{R}$. Then a curve $\xi : \mathbb{R} \rightarrow M$ is geodesic if and only if the projections $\rho(\xi) : \mathbb{R} \rightarrow M_0$ and $\mathbf{b}_\eta(\xi) : \mathbb{R} \rightarrow \mathbb{R}$ are geodesic.

Proof. (i) We have already seen in Corollary 4.4 that φ_t is measure-preserving. Given any $v \in T_x M$, the isometric splitting of $(M, g_{\nabla \mathbf{b}_\eta}, m)$ (Proposition 4.3) shows that $V(t) := D\varphi_t(v)$ is a parallel vector field with respect to $g_{\nabla \mathbf{b}_\eta}$ along the geodesic $\sigma(t) = \varphi_t(x)$. Hence it follows from Lemma 2.3 that $D_{\dot{\sigma}} V = D_{\dot{\sigma}}^\circ V = D_{\dot{\sigma}}^{g_{\nabla \mathbf{b}_\eta}} V \equiv 0$. Thus we obtain

$$\frac{\partial}{\partial t} [F(V)^2] = \frac{\partial}{\partial t} [g_V(V, V)] = 2g_V(D_{\dot{\sigma}} V, V) \equiv 0$$

and φ_t is isometric (we were again indebted to the fine property $D_{\dot{\sigma}}^\circ V = D_{\dot{\sigma}}^\circ V$ of Berwald spaces).

(ii) The total convexity in Lemma 5.1 guarantees that $(M_0, F|_{TM_0})$ is of Berwald type (via the characterization (e) in [BCS, Theorem 10.2.1] for instance). In order to see $\text{Ric}_{N-1} \geq 0$, recall from Corollary 4.4 that m_0 enjoys $m = m_0 \times \mathbf{L}^1$. Fix a unit vector $v \in U_x M_0$ and extend it to a vector field V_0 on a neighborhood $U \subset M_0$ of x such that all integral curves of V_0 are geodesic. We further extend V_0 to V on $U \times (-\varepsilon, \varepsilon) \subset M$ by $V(y, t) = (V_0(y), 0) \in T_{(y,t)} M$. Then all integral curves of V are geodesic and we deduce

$$\text{Ric}_{N-1}^{M_0}(v) = \text{Ric}_N^M((v, 0)) \geq 0$$

directly from the definition of Ric_N (since $(N-1) - (n-1) = N-n$).

(iii) We can split the geodesic equation (2.7) for M into those for M_0 and \mathbb{R} by virtue of a special property of Berwald spaces. Take an open set $U \subset M_0$ and a coordinate (x^1, x^2, \dots, x^n) of $U \times \mathbb{R} \subset M$ such that $\mathbf{b}_\eta(x) = x^n$ and that $(\partial/\partial x^i)_{i=1}^n$ is orthogonal with respect to $g_{\nabla \mathbf{b}_\eta}$. As Γ_{jk}^i is constant on every tangent space $T_x M \setminus 0$, let us denote it by $\Gamma_{jk}^i(x)$. We shall calculate

$$\begin{aligned} \Gamma_{jk}^i(x) &= \Gamma_{jk}^i(\nabla \mathbf{b}_\eta(x)) \\ &= \gamma_{jk}^i(\nabla \mathbf{b}_\eta(x)) - \sum_{m=1}^n (A_{jim} N_k^m + A_{ikm} N_j^m - A_{jkm} N_i^m)(\nabla \mathbf{b}_\eta(x)). \end{aligned}$$

Note that

$$\frac{\partial g_{ij}}{\partial x^n}(\nabla \mathbf{b}_\eta) = 0 \quad \text{for } i, j = 1, 2, \dots, n$$

by (i), and that

$$\frac{\partial g_{in}}{\partial x^j}(\nabla \mathbf{b}_\eta) = 0 \quad \text{for } i, j = 1, 2, \dots, n$$

by $g_{\nabla \mathbf{b}_\eta}(\nabla \mathbf{b}_\eta, TM_0) = 0$ and $F(\nabla \mathbf{b}_\eta) \equiv 1$. Hence $\gamma_{jk}^i(\nabla \mathbf{b}_\eta(x)) = 0$ unless $i, j, k \neq n$. In particular, it follows from (2.4) that $N_j^i(\nabla \mathbf{b}_\eta(x)) = 0$ for all $i, j = 1, 2, \dots, n$. Therefore we have $\Gamma_{jk}^i(x) = 0$ unless $i, j, k \neq n$, so that

$$D_\xi \dot{\xi} = \sum_{i=1}^{n-1} \left\{ \ddot{\xi}^i + \sum_{j,k=1}^{n-1} \Gamma_{jk}^i(\xi) \dot{\xi}^j \dot{\xi}^k \right\} \frac{\partial}{\partial x^i} \Big|_\xi + \ddot{\xi}^n \frac{\partial}{\partial x^n} \Big|_\xi.$$

Thus the geodesic equation is split and we complete the proof. \square

We remark that, different from the Riemannian case, one can not reconstruct (M, F) only from $(M_0, F|_{TM_0})$. Indeed, given $x \in M_0$, all we know is $F|_{T_x M_0}$ and the fact that $T_x M_0 \perp \nabla \mathbf{b}_\eta(x)$ with respect to $g_{\nabla \mathbf{b}_\eta}$. They give us only a little information about $F|_{T_x M \setminus T_x M_0}$ (see the related discussion in Remark 5.5 below).

We can iterate the procedure in Proposition 5.2 and obtain the following.

Corollary 5.3 (i) *There exists a k -parameter family of measure-preserving isometries $\varphi_p : M \rightarrow M$, $p \in \mathbb{R}^k$, and an $(n - k)$ -dimensional totally convex, geodesically complete submanifold $M' \subset M$ such that*

- $(M', F|_{TM'})$ does not contain a straight line;
- $(M', F|_{TM'}, m')$ is of Berwald type with $\text{Ric}_{N-k} \geq 0$, where $m' := m|_{M'}$;
- $\bigsqcup_{p \in \mathbb{R}^k} \varphi_p(M') = M$;
- $\varphi_{p+q} = \varphi_q \circ \varphi_p$ for any $p, q \in \mathbb{R}^k$.

In particular, (M, m) admits a diffeomorphic, measure-preserving splitting $(M, m) = (M' \times \mathbb{R}^k, m' \times \mathbf{L}^k)$.

(ii) *For each $x \in M$, $\Sigma_x := \{\varphi_p(x)\}_{p \in \mathbb{R}^k}$ is a k -dimensional submanifold of M whose flag curvature (with respect to the restriction of F) vanishes everywhere.*

(iii) *For any $x, y \in M'$, $(\Sigma_x, F|_{T\Sigma_x})$ is isometric to $(\Sigma_y, F|_{T\Sigma_y})$.*

Proof. (i) There is nothing to prove if (M, F) does not contain a straight line ($k = 0$, $M' = M$). If (M, F) contains a straight line $\eta_1 : \mathbb{R} \rightarrow M$, then Proposition 5.2 provides a one-parameter family $\{\varphi_t^{(1)}\}_{t \in \mathbb{R}}$ of measure-preserving isometries and an $(n-1)$ -dimensional totally convex, geodesically complete submanifold $M_1 \subset M$ of $\text{Ric}_{N-1} \geq 0$. Suppose that M_1 contains a straight line $\eta_2 : \mathbb{R} \rightarrow M_1$ again. We remark that η_2 is a straight line also as a curve in M thanks to Proposition 5.2(iii). Similarly to $\varphi_t^{(1)}$, we obtain $\varphi_s^{(2)} : M_1 \rightarrow M_1$, $s \in \mathbb{R}$, and $M_2 \subset M_1$. Now, define $\varphi_{(t,s)} : M = M_1 \times \mathbb{R} \rightarrow M$ by

$$\varphi_{(t,s)}(x, r) := \varphi_t^{(1)}(\varphi_s^{(2)}(x), r) = (\varphi_s^{(2)}(x), r + t).$$

This map clearly preserves the measure m . To see that $\varphi_{(t,s)}$ is isometric, it is sufficient to show that the map $(x, r) \mapsto (\varphi_s^{(2)}(x), r)$ is isometric. Consider the vector field $V := (\nabla \mathbf{b}_{\eta_2}, 0) \in TM_1 \times T\mathbb{R} = TM$. Note that all integral curves of V are straight lines and mutually bi-asymptotic (compare two lines $\xi, \zeta \subset M_1$ and then ξ and $\varphi_t \circ \xi$). Therefore V coincides with the gradient vector field of the Busemann function on M for the line η_2 (see Lemma 4.2). Hence the argument as in Proposition 5.2(i) ensures that $\varphi_{(t,s)}$ is isometric. The commutativity $\varphi_{(t+t', s+s')} = \varphi_{(t,s)} \circ \varphi_{(t',s')}$ is straightforward from the construction. Iterating this procedure as far as possible provides the desired space M' containing no straight line as well as the family of isometries $\{\varphi_p\}_{p \in \mathbb{R}^k}$.

(ii) The flatness is a consequence of the rigidity theorem for Berwald surfaces (Theorem 2.11). As Σ_x is totally geodesic (in other words, *locally* totally convex) and hence of

Berwald type by Proposition 5.2(iii), we can apply Theorem 2.11 to each two-dimensional subspace Π of Σ_x (precisely, $\Pi = \{\varphi_p(x)\}_{p \in P}$ for a two-dimensional affine subspace $P \subset \mathbb{R}^k$). Proposition 5.2(iii) also verifies that Π must be flat even if it is Riemannian, therefore Σ_x is flat.

(iii) Take a minimal geodesic $\xi : [0, 1] \rightarrow M$ from x to y and $p \in \mathbb{R}^k$. By Theorem 4.6, the parallel transport along $\varphi_p \circ \xi$ sends $\nabla \mathbf{b}_\eta(\varphi_p(x))$ to $\nabla \mathbf{b}_\eta(\varphi_p(y))$ for any straight line $\eta(t) = \varphi_{tq}(x)$ for some $q \in \mathbb{R}^k$. Since parallel transports are isometric in Berwald spaces (Proposition 2.10), we conclude that Σ_x is isometric to Σ_y . \square

It is unclear from the infinitesimal discussion in Proposition 5.2 and Corollary 5.3 if Σ_x is *globally* totally convex in M (see also Remark 5.5 below). One may be able to split M' again in a slightly different way as follows.

Theorem 5.4 (Isometric splitting of Berwald spaces) *Suppose that (M, F, m) is of Berwald type and forward and backward complete, and satisfies $\text{Ric}_N \geq 0$ for some $N \in [n, \infty]$. Assume also that Ψ is bounded above if $N = \infty$. Then the following hold.*

(i) *There exists an l -parameter family of measure-preserving isometries $\varphi_p : M \rightarrow M$, $p \in \mathbb{R}^l$, and an $(n-l)$ -dimensional totally convex, geodesically complete submanifold $\overline{M} \subset M$ such that*

- *any Busemann function associated with a straight line in M is constant on \overline{M} ;*
- *$(\overline{M}, F|_{T\overline{M}}, \overline{m})$ is of Berwald type with $\text{Ric}_{N-l} \geq 0$, where $\overline{m} := m|_{\overline{M}}$;*
- *$\bigsqcup_{p \in \mathbb{R}^l} \varphi_p(\overline{M}) = M$;*
- *$\varphi_{p+q} = \varphi_q \circ \varphi_p$ for any $p, q \in \mathbb{R}^l$.*

In particular, (M, m) admits a diffeomorphic, measure-preserving splitting $(M, m) = (\overline{M} \times \mathbb{R}^l, \overline{m} \times \mathbf{L}^l)$.

(ii) *For each $x \in M$, $\Sigma_x := \{\varphi_p(x)\}_{p \in \mathbb{R}^l}$ is an l -dimensional submanifold of M whose flag curvature vanishes everywhere.*

(iii) *For any $x, y \in \overline{M}$, $(\Sigma_x, F|_{T\Sigma_x})$ is isometric to $(\Sigma_y, F|_{T\Sigma_y})$.*

Proof. (i) We start from the splitting $(M, m) = (M' \times \mathbb{R}^k, m' \times \mathbf{L}^k)$ in Corollary 5.3, and shall use the same notations. Assume that there is a straight line $\eta : \mathbb{R} \rightarrow M$ such that the Busemann function \mathbf{b}_η is not constant on M' . We put $M'' := \mathbf{b}_\eta^{-1}(0) \cap M'$ and first observe that it is a totally convex, geodesically complete, $(n-k-1)$ -dimensional submanifold of Berwald type with $\text{Ric}_{N-k-1} \geq 0$. We next split M' . In the canonical coordinate (x, r_1, \dots, r_k) of $M' \times \mathbb{R}^k$, $\dot{\eta}(0) \in T_{\eta(0)}M$ is written as (v, a_1, \dots, a_k) with nonzero v . As $\nabla \mathbf{b}_\eta$ is a parallel vector field (and so are $\nabla \mathbf{b}_{\eta_1}, \dots, \nabla \mathbf{b}_{\eta_k}$), the $T\mathbb{R}^k$ -component of $\nabla \mathbf{b}_\eta$ is always (a_1, \dots, a_k) . (Observe also that the TM' -component of $\nabla \mathbf{b}_\eta|_{M'}$ is a parallel vector field on M' due to the linearity of the covariant derivative.) Moreover, for any $p \in \mathbb{R}^k$ and $x \in M'$, $\nabla \mathbf{b}_\eta(x)$ and $\nabla \mathbf{b}_\eta(\varphi_p(x))$ have the same $T_x M'$ -components since $\varphi_p(\exp_x(t\nabla \mathbf{b}_\eta))$ is bi-asymptotic to $\exp_x(t\nabla \mathbf{b}_\eta)$ (see Lemma 4.2). Therefore the one-parameter family of measure-preserving isometries $\psi_t : M \rightarrow M$, $t \in \mathbb{R}$, generated from $\nabla \mathbf{b}_\eta$ splits into

$$\psi_t = (\psi_t^{(1)}, \psi_t^{(2)}) \quad \text{with} \quad \psi_t^{(1)} : M' \rightarrow M', \quad \psi_t^{(2)} : \mathbb{R}^k \rightarrow \mathbb{R}^k,$$

and is written in the above coordinate as

$$\psi_t(x, r_1, \dots, r_k) = (\psi_t^{(1)}(x), r_1 + ta_1, \dots, r_k + ta_k).$$

Therefore $\psi_t^{(1)}$ derives a diffeomorphic, measure-preserving splitting $(M', m') = (M'' \times \mathbb{R}, m'|_{M'' \times \mathbf{L}^1})$. Iterating this construction as far as possible, we eventually obtain the desired submanifold \overline{M} .

(ii) and (iii) are shown in the same way as Corollary 5.3(ii), (iii). \square

Due to (i) above, starting from a point $x \in M$, we can characterize $\overline{M} \ni x$ by $\overline{M} = \bigcap_{\eta} \mathbf{b}_{\eta}^{-1}(0)$, where η runs over all straight lines parametrized as $\mathbf{b}_{\eta}(x) = 0$.

Remark 5.5 Obviously $k \leq l$ holds for k in Corollary 5.3 and l in Theorem 5.4. The author does not know any example satisfying $k < l$. We can verify $k = l$ if, for η in the proof of Theorem 5.4(i), the projection $\rho(\eta) : \mathbb{R} \rightarrow M'$ is a straight line (note the difference between $\eta_2 \subset M_1$ in Corollary 5.3(i) and $\eta \not\subset M'$ in Theorem 5.4(i)). The straightness is clear in the Riemannian case by the isometry of the splitting, however, unclear in the Finsler setting from our infinitesimal argument (we only know that $\rho(\eta)$ is geodesic by Proposition 5.2(iii)). The difficulty comes from the fact that a (Minkowski) normed space $(\mathbb{R}^n, |\cdot|)$ may contain a triple $v, w_1, w_2 \in \mathbb{R}^k \setminus \{0\}$ such that

$$g_v(v, w_1) = g_v(v, w_2) = 0, \quad |w_1| < |w_2|, \quad |w_1 + v| > |w_2 + v|.$$

Having Theorem 5.4 at hand, we can obtain some topological information of compact Berwald spaces along the lines of Cheeger and Gromoll's classical theory ([CG1], [CG2]).

Theorem 5.6 (A Betti number estimate) *Let (M, F, m) be a compact Finsler manifold of Berwald type satisfying $\text{Ric}_N \geq 0$ for some $N \in [n, \infty]$. Assume also that Ψ is bounded above if $N = \infty$. Denote by $(\widetilde{M}, \widetilde{F})$ the universal Finsler covering of (M, F) and by $\widetilde{M} = \overline{M} \times \mathbb{R}^l$ its splitting obtained in Theorem 5.4. Then we have the following.*

- (i) \overline{M} is compact.
- (ii) There exists a finite normal subgroup Ξ of the fundamental group $\pi_1(M)$ such that $\pi_1(M)/\Xi$ contains \mathbb{Z}^l as a normal subgroup of finite index.
- (iii) The first Betti number of M is at most l .

Proof. Let us first give a remark on the action of the isometry group $\text{Isom}(\widetilde{M}, \widetilde{F})$. For each $\Phi \in \text{Isom}(\widetilde{M}, \widetilde{F})$ and any straight line $\eta : \mathbb{R} \rightarrow M$, it clearly holds $\mathbf{b}_{\Phi \circ \eta} \circ \Phi = \mathbf{b}_{\eta}$. Hence $\Phi(\overline{M})$ inherits the property that all Busemann functions are constant.

(i) Let $\Gamma \subset \text{Isom}(\widetilde{M}, \widetilde{F})$ be the deck transformation group of the universal covering $\pi : \widetilde{M} \rightarrow M$. Since \overline{M} is compact, we can take a compact fundamental domain $K \subset \overline{M}$ such that $\Gamma(K) = \widetilde{M}$. Now, if \overline{M} is not compact, then there is a ray $\eta : [0, \infty) \rightarrow M$ contained in \overline{M} . For each $i \in \mathbb{N}$, we can choose $\Phi^{(i)} \in \Gamma$ satisfying $\Phi^{(i)}(\eta(i)) \in K$. Then $\eta_i(t) := \Phi^{(i)}(\eta(i+t))$, $t \in [-i, \infty)$, is a globally minimizing geodesic with $\eta_i(0) \in K$, and the compactness of K ensures that $\{\eta_i\}_{i \in \mathbb{N}}$ has a subsequence convergent to some

geodesic $\eta_\infty : \mathbb{R} \rightarrow M$ ($\dot{\eta}_i(0) \rightarrow \dot{\eta}_\infty(0)$ to be precise). This geodesic η_∞ is clearly a straight line, and the Busemann function \mathbf{b}_{η_∞} is not constant on $\Phi^{(i)}(\overline{M})$ for large i (since $\dot{\eta}_i \in T[\Phi^{(i)}(\overline{M})]$). This is a contradiction, so that \overline{M} is compact.

(ii) Fix $\Phi \in \mathbf{Isom}(\widetilde{M}, \widetilde{F})$. On the one hand, for each $p \in \mathbb{R}^l$, $\Phi(\varphi_p(\overline{M}))$ coincides with $\varphi_q(\overline{M})$ for some $q \in \mathbb{R}^l$ from the construction of \overline{M} (since Φ preserves Busemann functions). On the other hand, for each $x \in \overline{M}$, we find that $\Phi(\Sigma_x) = \Sigma_y$ for some $y \in \overline{M}$ thanks to the compactness of \overline{M} . Then the projection

$$\Phi_1 := \varphi_{p-q} \circ \Phi|_{\varphi_p(\overline{M})} \in \mathbf{Isom}(\varphi_p(\overline{M}), \widetilde{F}) = \mathbf{Isom}(\overline{M}, \widetilde{F})$$

(we identified $\mathbf{Isom}(\varphi_p(\overline{M}), \widetilde{F})$ with $\mathbf{Isom}(\overline{M}, \widetilde{F})$ via φ_p) is independent of the choice of $p \in \mathbb{R}^l$ because Φ preserves the \mathbb{R}^l -directions (that is, $\Phi(\Sigma_x) = \Sigma_y$ and $\Phi_1(x) = y$). We similarly see that $\Phi_2 := \Phi|_{\Sigma_x} \in \mathbf{Isom}(\Sigma_x, \widetilde{F})$ is independent of the choice of $x \in \overline{M}$ by identifying Σ_x and Σ_y via the coordinate of \mathbb{R}^l (Σ_x and Σ_y are actually isometric by Theorem 5.4(iii)). Consequently, we obtain

$$\mathbf{Isom}(\widetilde{M}, \widetilde{F}) = \mathbf{Isom}(\overline{M}, \widetilde{F}) \times \mathbf{Isom}(\Sigma_x, \widetilde{F}).$$

Since $\mathbf{Isom}(\overline{M}, \widetilde{F})$ is compact, $\Xi := \{\Phi \in \Gamma \mid \Phi_2 = \text{Id}_{\Sigma_x}\}$ is a finite normal subgroup of $\Gamma = \pi_1(M)$. The covering

$$\widetilde{M} = \overline{M} \times \mathbb{R}^l \rightarrow \widetilde{M}/\Xi =: \overline{M}' \times \mathbb{R}^l$$

leads

$$\mathbf{Isom}(\widetilde{M}/\Xi, \widetilde{F}) = \mathbf{Isom}(\overline{M}', \widetilde{F}) \times \mathbf{Isom}(\Sigma_x, \widetilde{F}) \supset \Gamma/\Xi,$$

and the projection of Γ/Ξ to $\mathbf{Isom}(\Sigma_x, \widetilde{F})$ is an isomorphism into its image. As Σ_x is flat (Theorem 5.4(ii)), $\mathbf{Isom}(\Sigma_x, \widetilde{F})$ is a subgroup of the isometry group $\mathbf{Isom}(\mathbb{R}^l)$ of \mathbb{R}^l with the standard Euclidean metric. Thus Γ/Ξ is isomorphic to a discrete uniform subgroup of $\mathbf{Isom}(\mathbb{R}^l)$. Therefore Γ/Ξ is isomorphic to a crystallographic group of \mathbb{R}^l , and contains \mathbb{Z}^l as a normal subgroup of finite index by the Bieberbach theorem.

(iii) This is a consequence of (ii). □

Our argument was indebted to various fine properties of Berwald spaces, even at the early stage (see Lemma 5.1, Proposition 5.2). It is unclear if any of results in this section can be generalized to general Finsler manifolds.

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